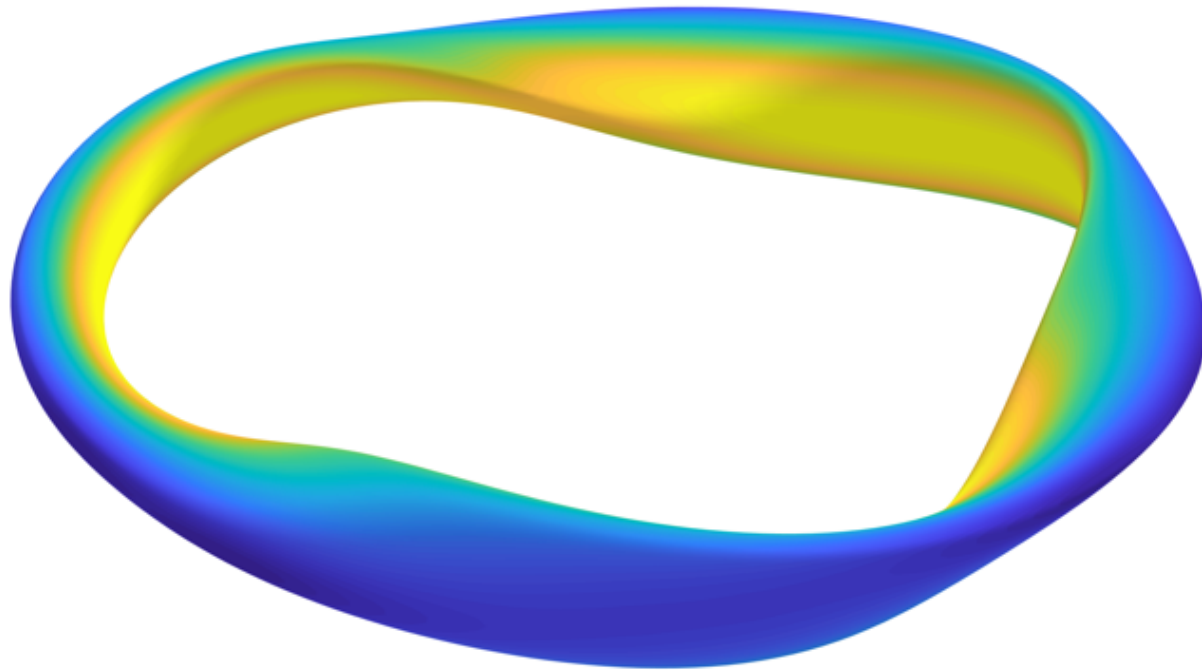


# Optimized stellarators without optimization:

*Direct construction of quasisymmetric shapes*



Matt Landreman (Maryland), Wrick Sengupta (NYU), Gabe Plunk (IPP-Greifswald)

[arXiv:1809.10246](https://arxiv.org/abs/1809.10246)

<https://github.com/landreman/quasisymmetry>

- In quasisymmetric designs to date, optimization has been done using “textbook” optimization algorithms to minimize symmetry-breaking Fourier modes in  $B$ .
  - Many local minima, so result depends on initial guess.
  - Never sure you’ve found all the interesting regions of parameter space.
  - Little insight as to the amount of freedom in the solution.

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  - Little insight as to the amount of freedom in the solution.
- For a complementary approach without these shortcomings, here we extend Garren & Boozer (1991).
  - Usually cited as a proof that quasisymmetry cannot be achieved to  $(a/R)^3$ .
  - Less well known that it contains a useful constructive procedure.
  - Provides insight & initial conditions for stellopt.

# Main idea of Garren-Boozer: expand position vector using Frenet frame, equate 2 representations of $\mathbf{B}$ .

$$\text{Frenet frame } (\mathbf{t}, \mathbf{n}, \mathbf{b}): \quad \frac{\partial \mathbf{r}_0}{\partial \ell} = \mathbf{t}, \quad \frac{d\mathbf{t}}{d\ell} = \kappa \mathbf{n}, \quad \frac{d\mathbf{n}}{d\ell} = -\kappa \mathbf{t} + \tau \mathbf{b}, \quad \frac{d\mathbf{b}}{d\ell} = -\tau \mathbf{n}$$

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$$\text{Dual relations: } \nabla r = \left[ \frac{\partial \mathbf{r}}{\partial r} \cdot \frac{\partial \mathbf{r}}{\partial \theta} \times \frac{\partial \mathbf{r}}{\partial \zeta} \right]^{-1} \frac{\partial \mathbf{r}}{\partial \theta} \times \frac{\partial \mathbf{r}}{\partial \zeta}, \quad \text{cyclic permutations.}$$

# Garren & Boozer derived a 1D “Ricatti” ODE for quasisymmetry to $O(a/R)$ .

$$\frac{d\sigma}{d\zeta} + \iota \left[ \frac{\bar{\eta}^4}{\kappa^4} + 1 + \sigma^2 \right] - 2 \frac{\bar{\eta}^2}{\kappa^2} [I_2 - \tau] = 0$$

$\kappa(\zeta)$  = curvature,     $\tau(\zeta)$  = torsion,     $I_2$  = current density

$\iota$  = rotational transform,     $\bar{\eta}$  = some constant

$$\mathbf{r}(r, \theta, \zeta) = \mathbf{r}_0(\zeta) + r X_{1c}(\zeta) \cos \theta \mathbf{n}(\zeta) + r [Y_{1s}(\zeta) \sin \theta + Y_{1c}(\zeta) \cos \theta] \mathbf{b}(\zeta) + O(r^2)$$

$$X_{1c}(\zeta) = \bar{\eta} / \kappa(\zeta)$$

$$Y_{1s}(\zeta) = \kappa(\zeta) / \bar{\eta}$$

$$Y_{1c}(\zeta) = \sigma(\zeta) \kappa(\zeta) / \bar{\eta}$$

**We have a new proof of existence & uniqueness of solutions, even though the GB ODE is nonlinear.**

Given  $P(\zeta) > 0$ ,  $Q(\zeta)$ , and  $\sigma(0)$ , with  $P(\zeta)$  and  $Q(\zeta)$

$2\pi$ -periodic, bounded, and integrable, a solution to

$$\frac{d\sigma}{d\zeta} + \iota(P + \sigma^2) + Q = 0 \quad (1)$$

is a pair  $\{\iota, \sigma(\zeta)\}$  solving (1) where  $\sigma(\zeta)$  is  $2\pi$ -periodic.

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**Theorem:** A solution exists and it is unique.

$\Rightarrow$  Numerical solution is very robust.

# **We can now precisely state the amount of freedom in 1<sup>st</sup>-order-in-r quasisymmetry.**

- For every magnetic axis shape (2 functions of  $\phi$ ) with nonvanishing curvature, and 3 real numbers ( $\bar{\eta}$ ,  $\sigma(0)$ , and  $I_2$ ), there is precisely 1 way to shape the near-axis surfaces consistent with quasisymmetry.
  - For stellarator symmetry,  $\sigma(0) = 0$ .
  - In the usual case of no current density on axis, then  $I_2 = 0$ .

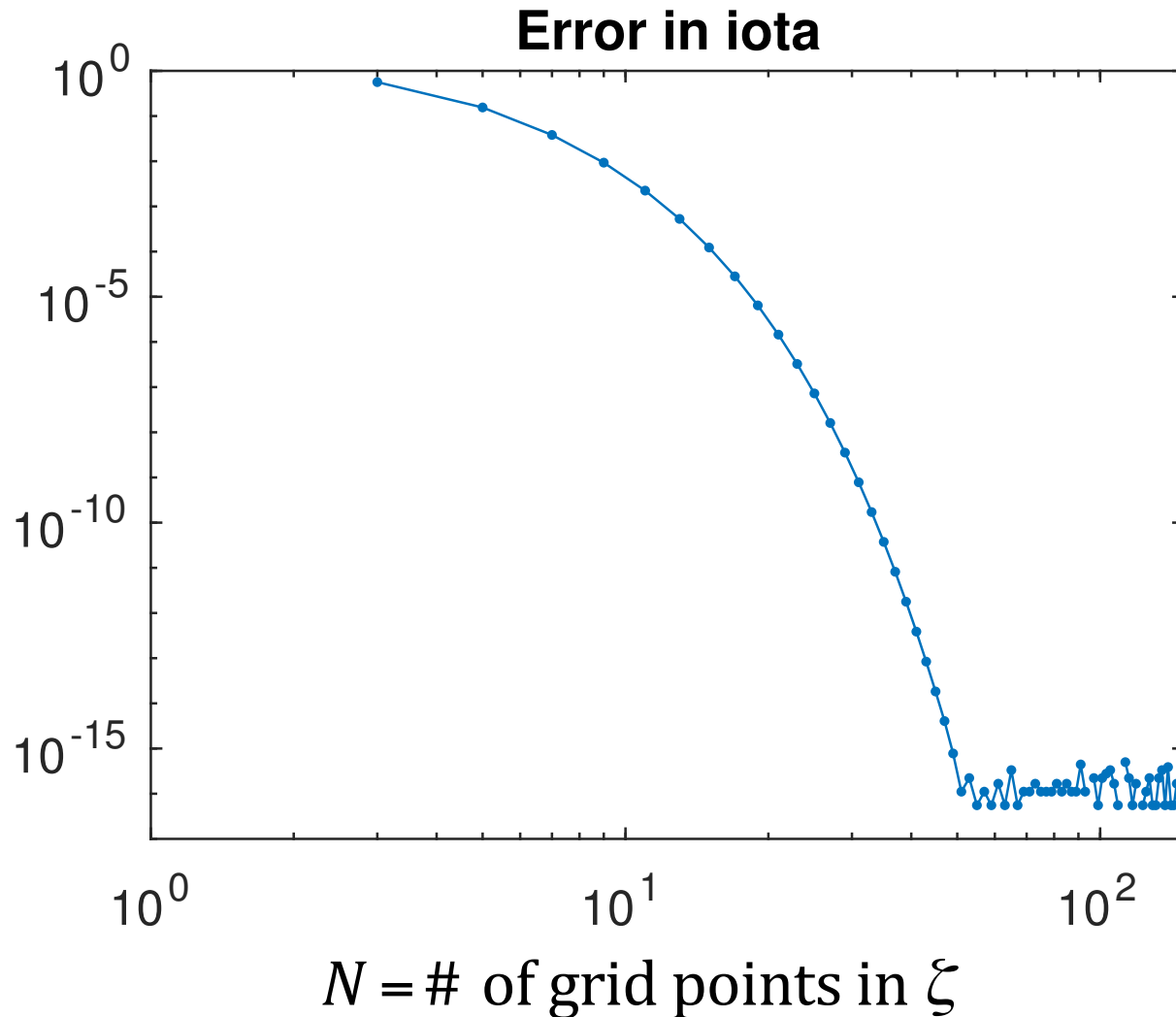
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- This solution may be quasi-axisymmetric or quasi-helically symmetric.
- However many of these solutions have absurdly high elongation.

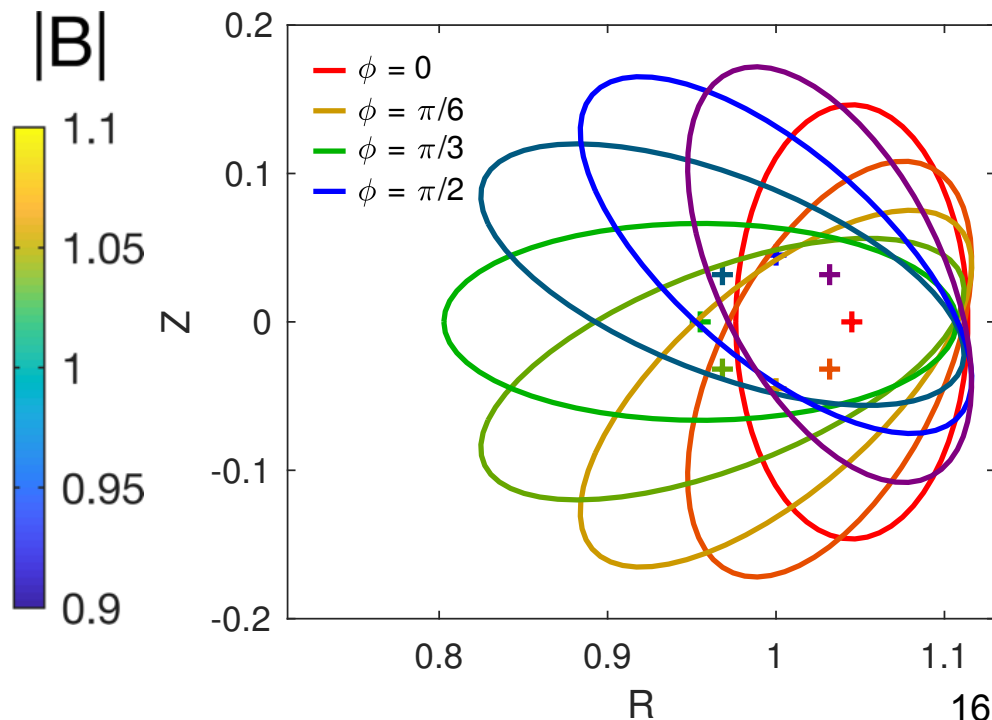
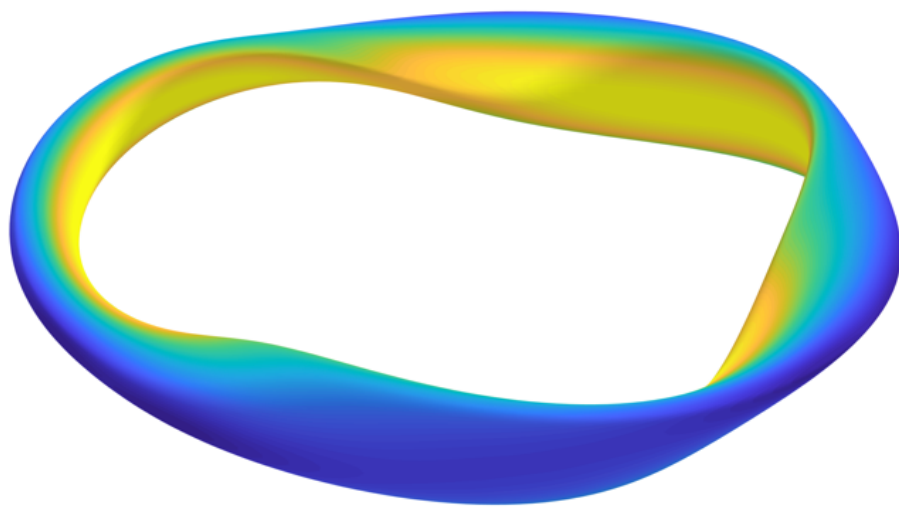
**The ODE is solved with spectral accuracy using  
pseudospectral discretization + Newton iteration**



# Quasi-axisymmetric example

**Inputs:** axis shape  $R_0(\phi) = 1 + 0.045 \cos(3\phi)$ ,  
 $Z_0(\phi) = -0.045 \sin(3\phi)$ ,  
 $\bar{\eta} = -0.9$ ,  $\sigma(0) = 0$

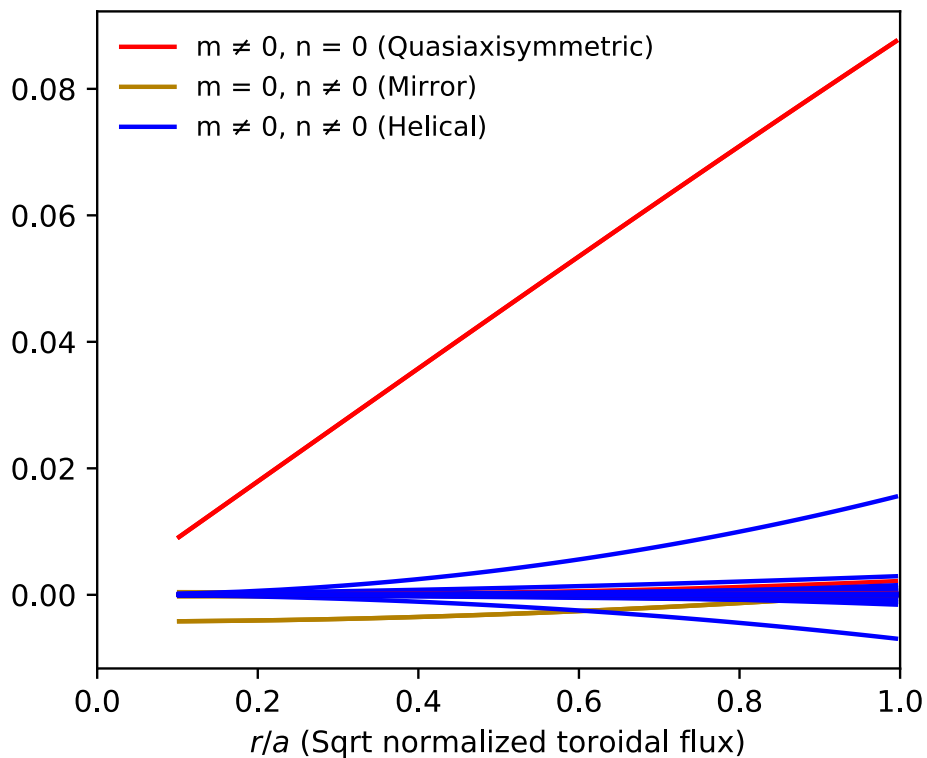
**Results:** ( $R/a = 10$ )



# Quasi-axisymmetric example

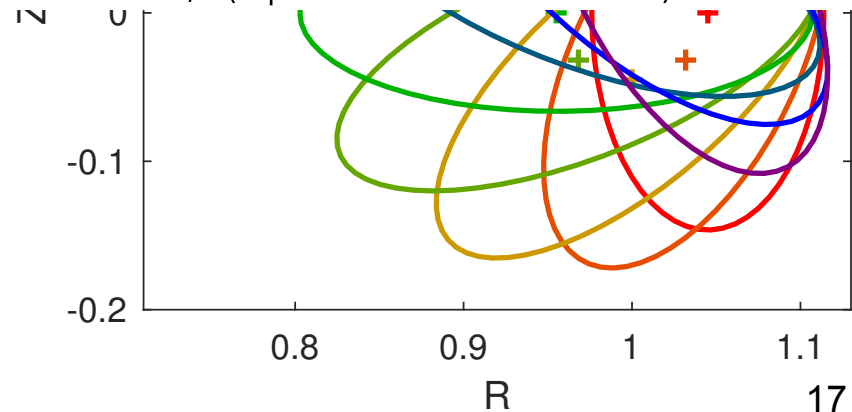
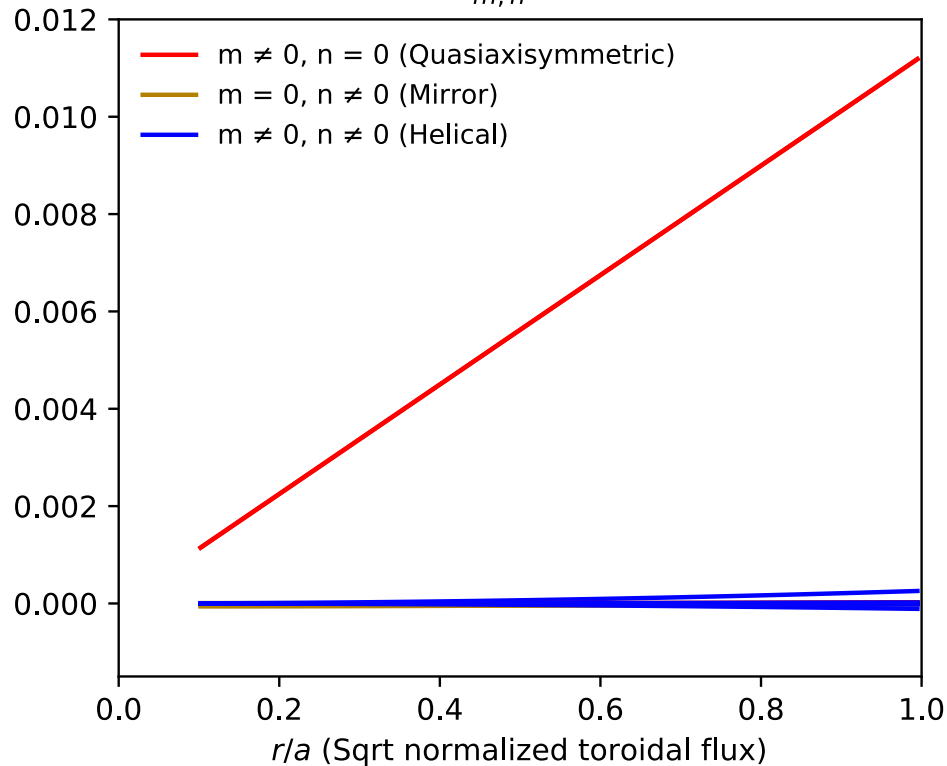
Aspect ratio 10

Fourier harmonics  $B_{m,n}$  in Boozer coordinates

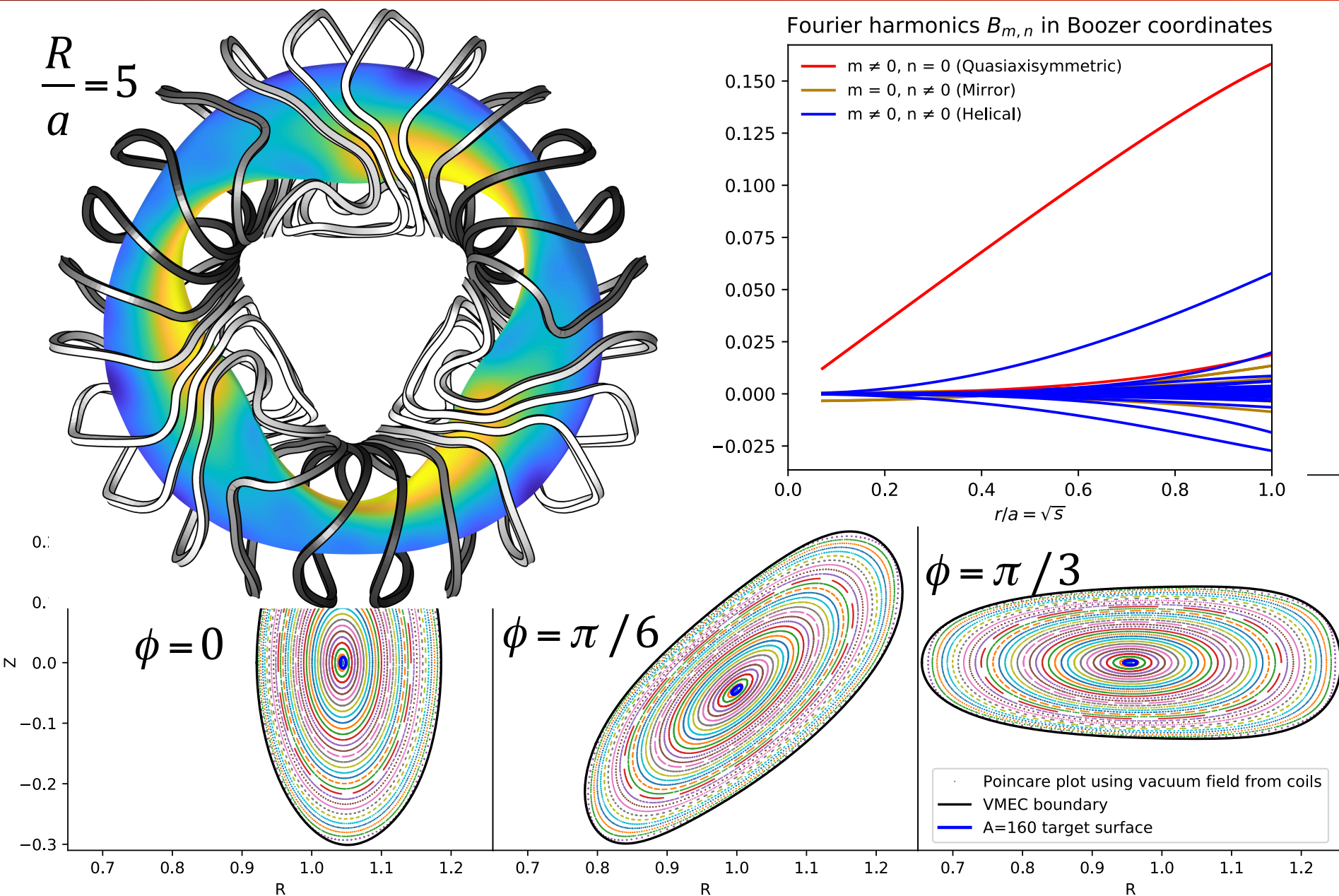


Aspect ratio 80

Fourier harmonics  $B_{m,n}$  in Boozer coordinates



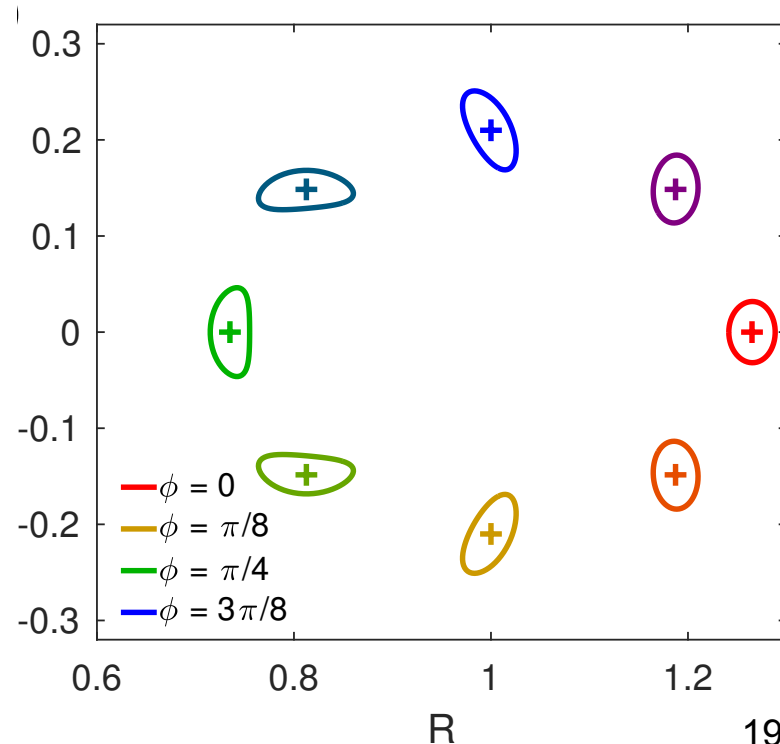
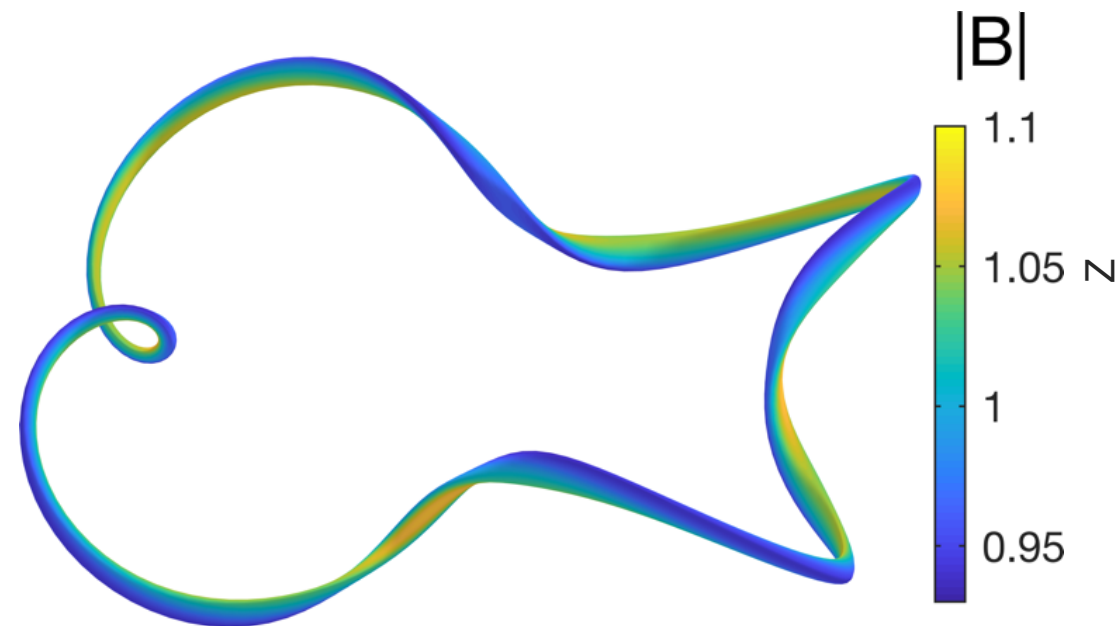
# If you generate coils to make a high-A surface, you actually often get a lower-A stellarator.



# Quasi-helical symmetry example

**Inputs:** axis shape  $R_0(\phi) = 1 + 0.265 \cos(4\phi)$ ,  
 $Z_0(\phi) = -0.21 \sin(4\phi)$ ,  
 $\bar{\eta} = -2.25$ ,  $\sigma(0) = 0$

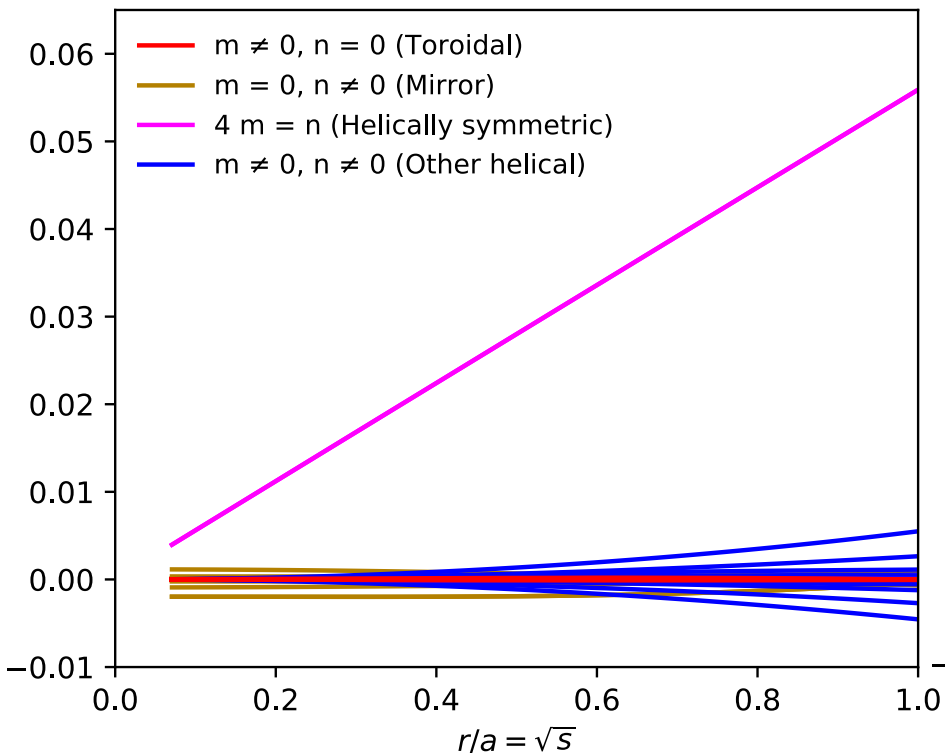
**Results:** ( $R/a = 40$ )



# Quasi-helical symmetry example

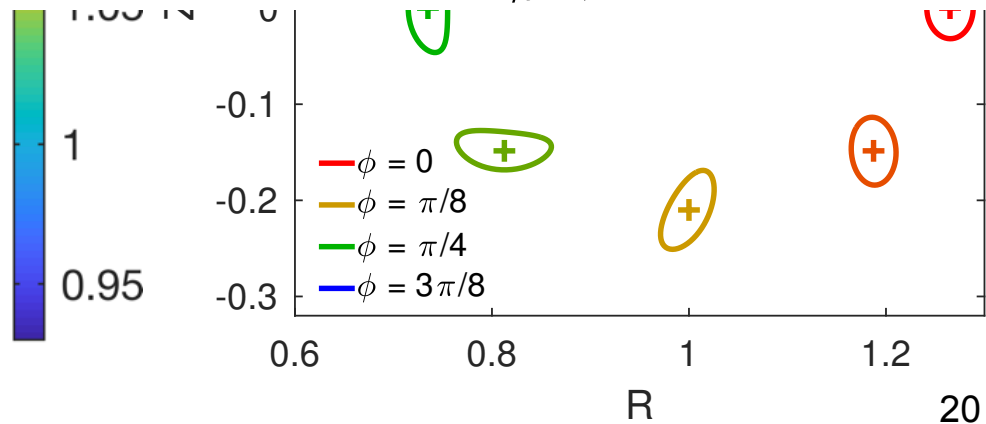
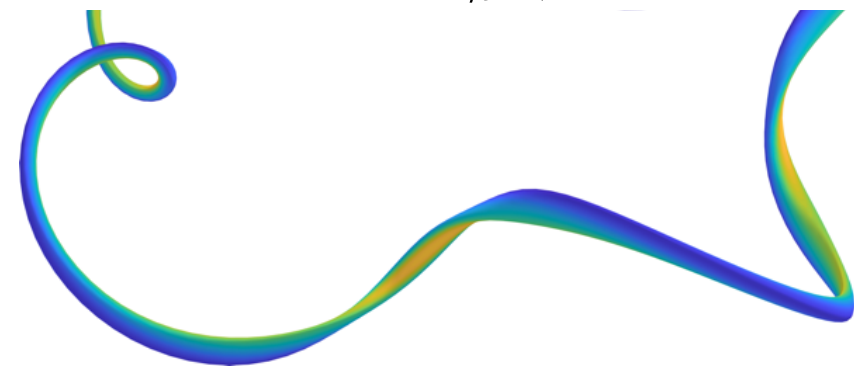
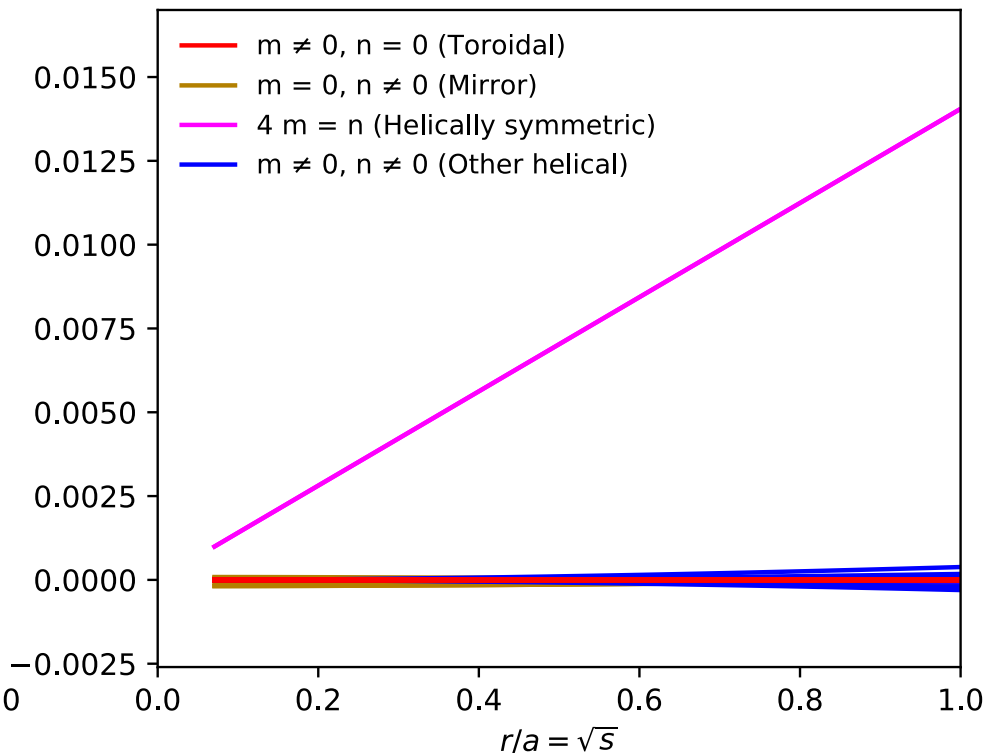
Aspect ratio 40

Fourier harmonics  $B_{m,n}$  in Boozer coordinates



Aspect ratio 160

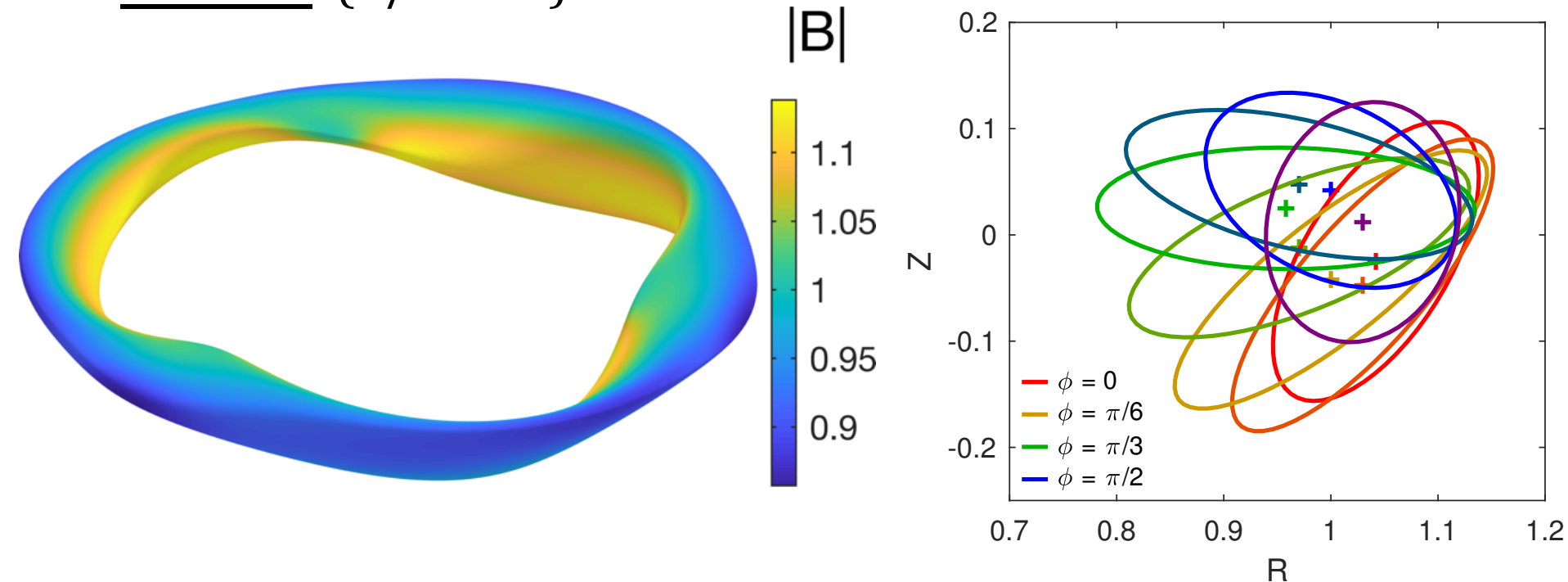
Fourier harmonics  $B_{m,n}$  in Boozer coordinates



# Idea from Greg Hammett: A quasi-axisymmetric stellarator without stellarator symmetry.

**Inputs:** axis shape  $R_0(\phi) = 1 + 0.042\cos(3\phi)$ ,  
 $Z_0(\phi) = -0.042\sin(3\phi) - 0.025\cos(3\phi)$ ,  
 $\bar{\eta} = -1.1$ ,  $\sigma(0) = -0.6$

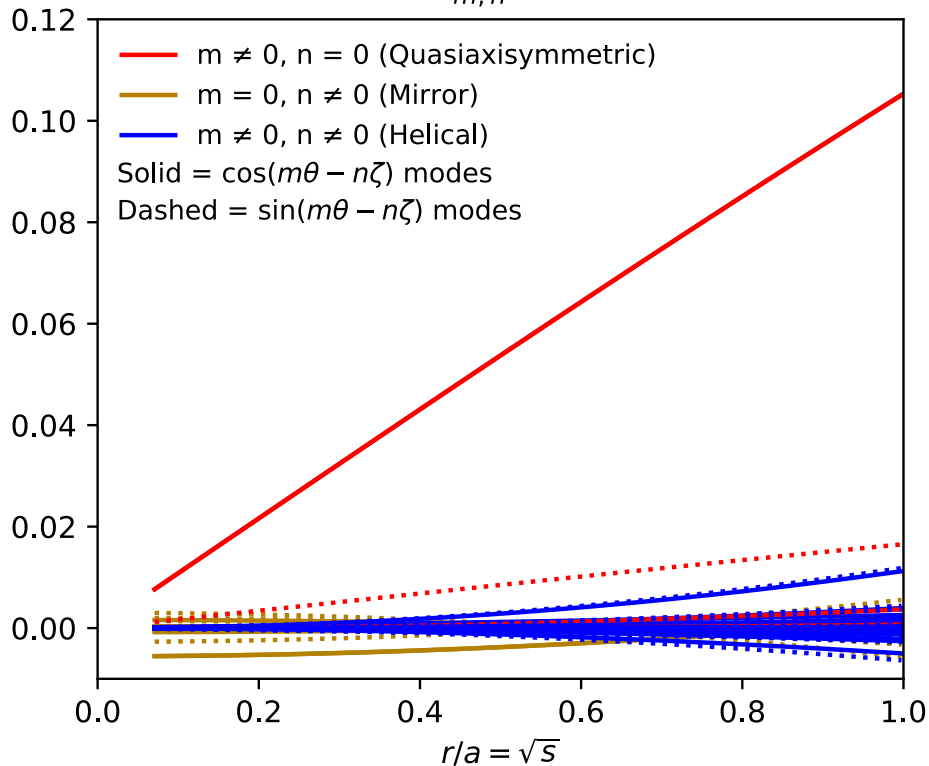
**Results:** ( $R/a = 10$ )



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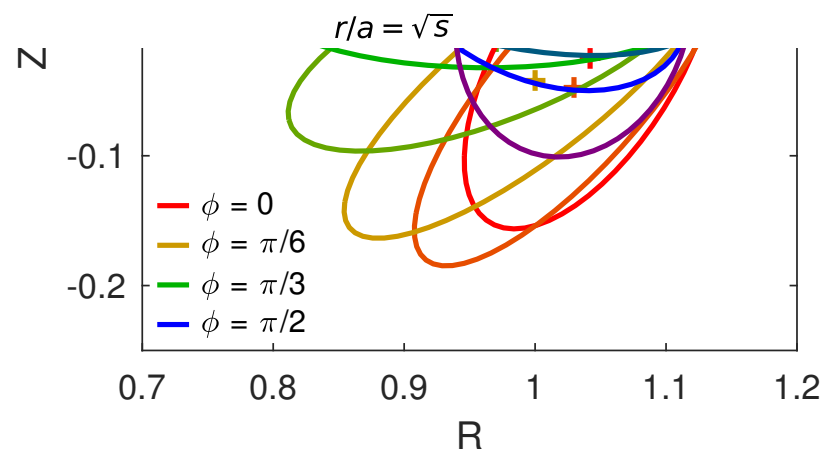
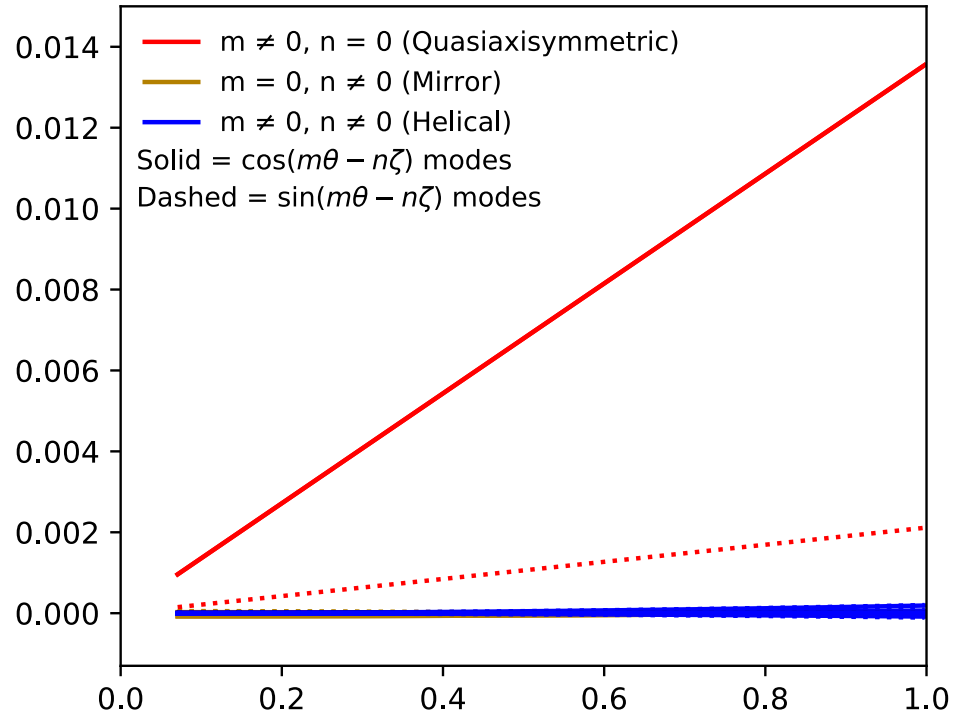
Aspect ratio 10

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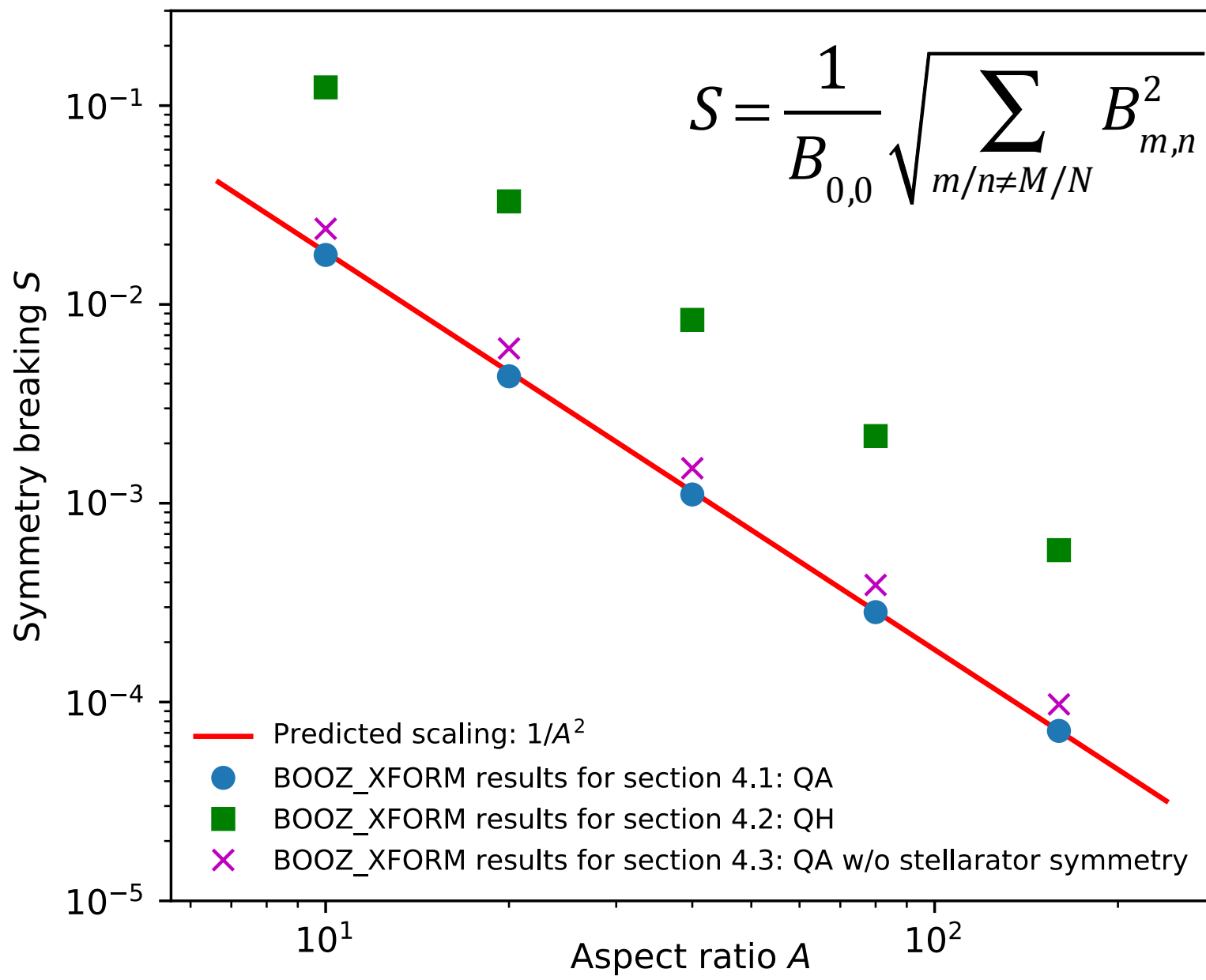


Aspect ratio 80

Fourier harmonics  $B_{m,n}$  in Boozer coordinates

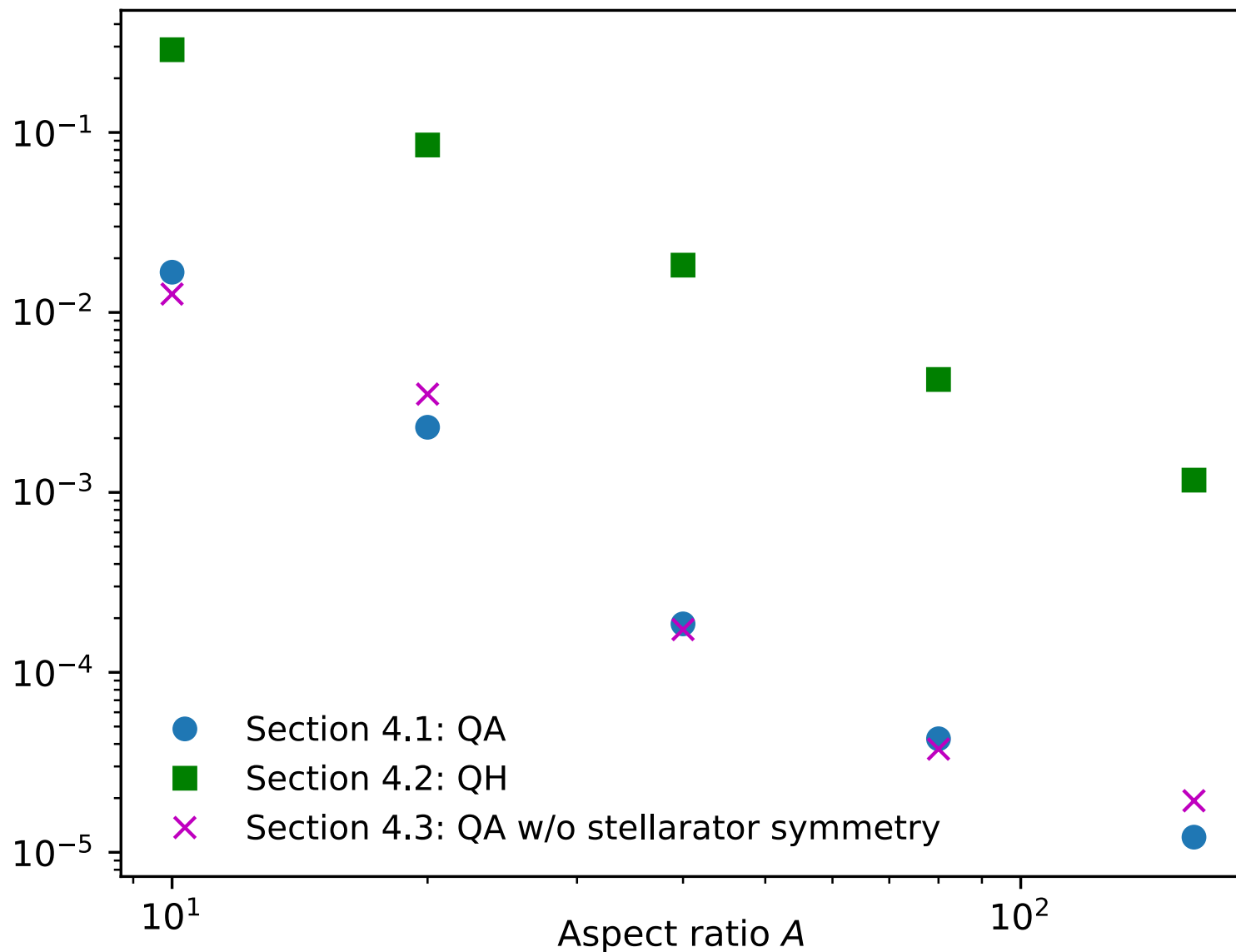


# The symmetry-breaking Fourier amplitudes scale as predicted.



# The rotational transform computed by VMEC converges to the value computed by the Garren-Boozer approach.

Difference in rotational transform  $\iota$  between VMEC vs ODE



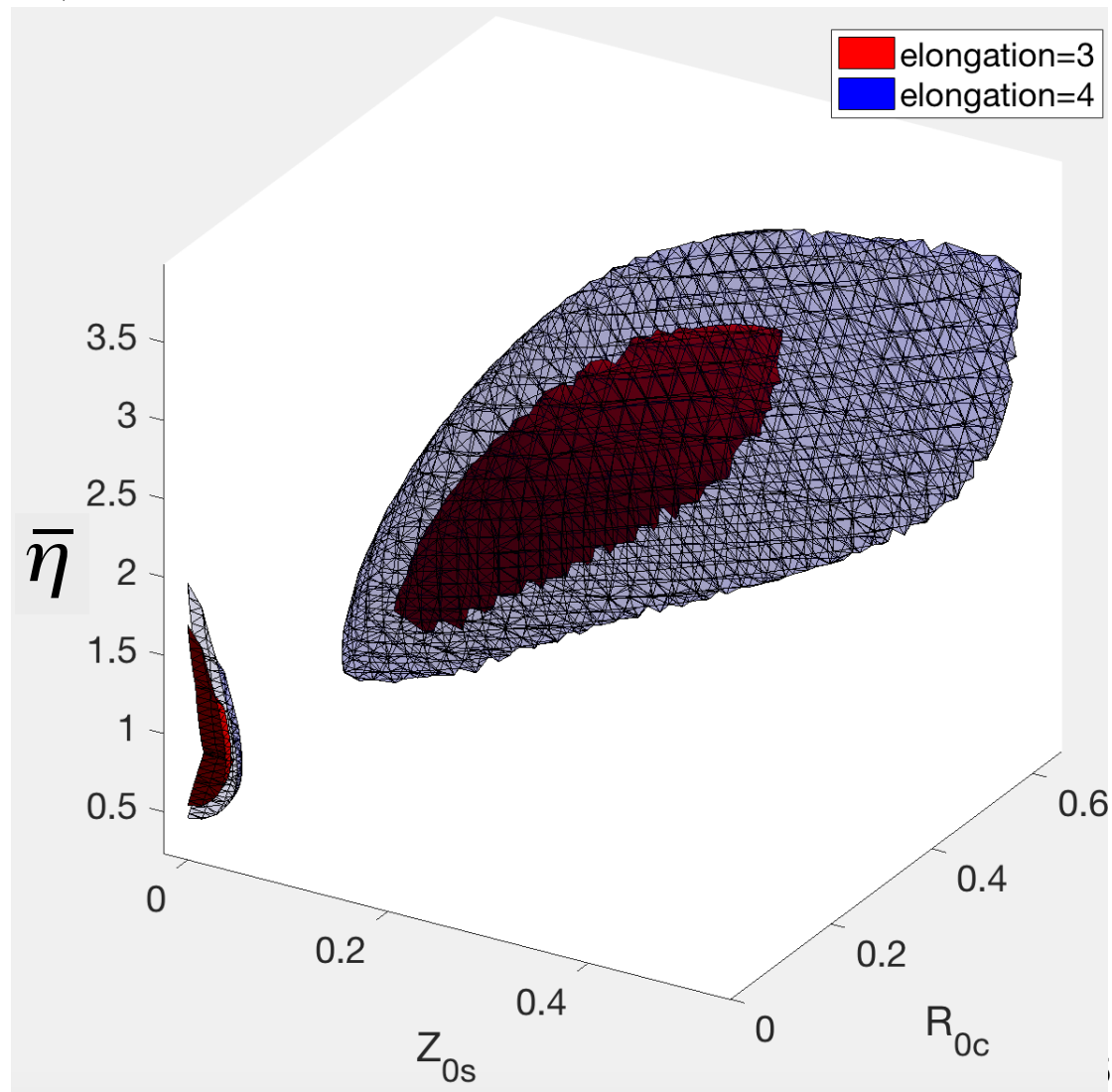
# This method enables fast scans over parameter space.

E.g. Scan over  $(R_{0c}, Z_{0s}, \bar{\eta})$  where magnetic axis shape is

$$R_0(\phi) = 1 + R_{0c} \cos(4\phi)$$

$$Z_0(\phi) = Z_{1s} \sin(4\phi)$$

274,560 solutions  
generated in <30s on  
my laptop.



Conclusion: The Garren-Boozer construction provides insight into the space of quasisymmetric shapes, & useful initial conditions for stellopt.

There are many extensions to pursue:

- Fully map the landscape of possible 1<sup>st</sup>-order quasisymmetric shapes by considering more Fourier modes in the axis shape. (How do I plot this?)
- Omnigenity.
- Can we construct shapes with quasisymmetry imposed at a mid-radius surface?
- 2<sup>nd</sup> order in distance from the axis.
- Connect to analysis of the difficulty of producing various plasma shapes.

**Extra slides**

# **The ODE is solved with spectral accuracy using pseudospectral discretization + Newton iteration**

Uniform grid in  $\phi$  with  $N$  points:  $\phi_1 = 0$ ,  $\phi_2 = 2\pi / (Nn_{fp})$ , ...,  $\phi_N = 2\pi(N-1) / (Nn_{fp})$ .

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Vector of  $N$  unknowns:  $\left( \iota, \sigma(\phi_2), \sigma(\phi_3), \dots, \sigma(\phi_N) \right)^T$

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